

Review of last lecture

1.  $\epsilon'_\lambda, \epsilon_\lambda, \epsilon', \epsilon$

2.  $\alpha'_\lambda, \alpha_\lambda, \alpha', \alpha$

3.  $\rho''_\lambda, \rho^{\lambda A}, \rho^{\lambda A'}, \rho^{\lambda A''}, \rho^{\lambda A''}$

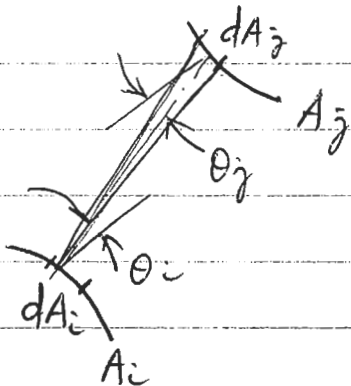
diffuse-gray  
 surface  
 emitter  
 reflector

4. Energy balance:  $\rho + \alpha + \tau = 1$

$\rho^{\lambda A} + \alpha'_\lambda + \rho^{\lambda A} = 1$

5. Kirchoff's law  $\alpha'_\lambda = \epsilon'_\lambda$

6.



diffuse-gray surfaces

$F_{dA_i-dA_j} = \frac{\cos\theta_i \cos\theta_j}{\pi r^2} dA_j$

=  $\frac{\text{power intercepted by } dA_j}{\text{total power leaving } dA_i}$

$$dF_{dA_i - dA_j} = \frac{\cos \theta_i \cos \theta_j}{\pi S^2} dA_j$$

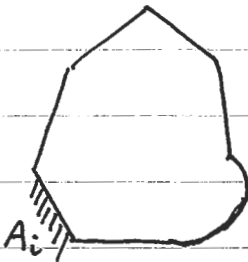
$$\Rightarrow dA_i dF_{dA_i - dA_j} = dA_j \cdot dF_{dA_j - dA_i}$$

law of reciprocity.

$$F_{dA_i - A_j} = \int_{A_j} \frac{\cos \theta_i \cos \theta_j}{\pi S^2} dA_j$$

$$F_{A_i - A_j} = \frac{1}{A_i} \int_{A_i} \int_{A_j} \frac{\cos \theta_i \cos \theta_j}{\pi S^2} dA_i dA_j$$

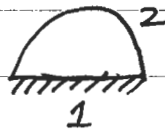
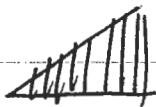
$$A_i F_{A_i - A_j} = A_j F_{A_j - A_i}$$



$$\sum_{i=1}^n A_i F_{A_i - A_j} = 1 \Rightarrow \sum_{j=1}^n F_{A_i - A_j} = 1$$

summation rule.

• Evaluation of View factors  
Simple cases



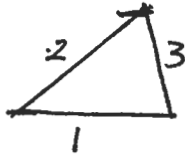
$$A_1 F_{12} = 1$$

~~$$A_2 F_{21} = A_1 F_{12}$$~~

$$A_2 F_{21} = A_1 F_{12}$$

$$F_{21} = \frac{A_1}{A_2}$$

$$F_{22} = 1 - F_{21} = 1 - \frac{A_1}{A_2}$$



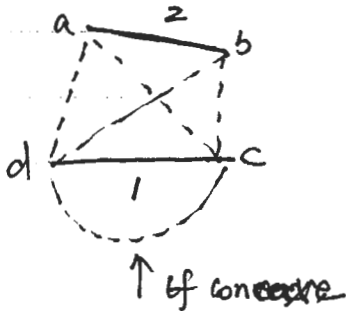
$$F_{12} + F_{13} = 1, \quad A_1 F_{12} = A_2 F_{21}$$

$$F_{21} + F_{23} = 1, \quad A_2 F_{23} = A_3 F_{32}$$

$$F_{31} + F_{32} = 1, \quad A_1 F_{13} = A_3 F_{31}$$

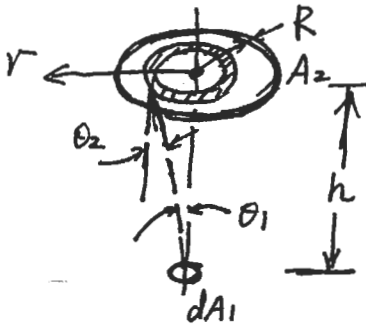
Solve :

$$F_{12} = \frac{A_1 + A_2 - A_3}{2A_1}$$



$$F_{12} = \frac{(ac+bd) - (ad+bc)}{2A_1} \quad \text{Hotter string rule.}$$

Direct Integration :



$$F_{12} = \int_{A_2} \frac{\cos\theta_1 \cos\theta_2}{\pi \delta^2} dA_2$$

$$dA_2 = 2\pi r dr$$

$$\cos\theta_1 = \cos\theta_2 = \frac{h}{\sqrt{h^2 + r^2}}$$

$$F_{12} = \int_0^R \frac{h^2}{\pi (h^2 + r^2)^2} 2\pi r dr$$

$$= h^2 \int_0^R \frac{dr^2}{(h^2 + r^2)^2}$$

$$= h^2 \left( -\frac{1}{h^2 + r^2} \right)_0^R = h^2 \left( \frac{1}{h^2} - \frac{1}{h^2 + R^2} \right)$$

$$= \frac{R^2}{h^2 + R^2}$$

Table  
Numerical

• ~~Radiosity~~ ~~and irradiation~~ ~~between~~ Radiative Heat Transfer Between Black Surfaces



$$Q_{1 \rightarrow 2} = F_{12} A_1 E_{b1}(T_1)$$

$$Q_{2 \rightarrow 1} = F_{21} A_2 E_{b2}(T_2)$$

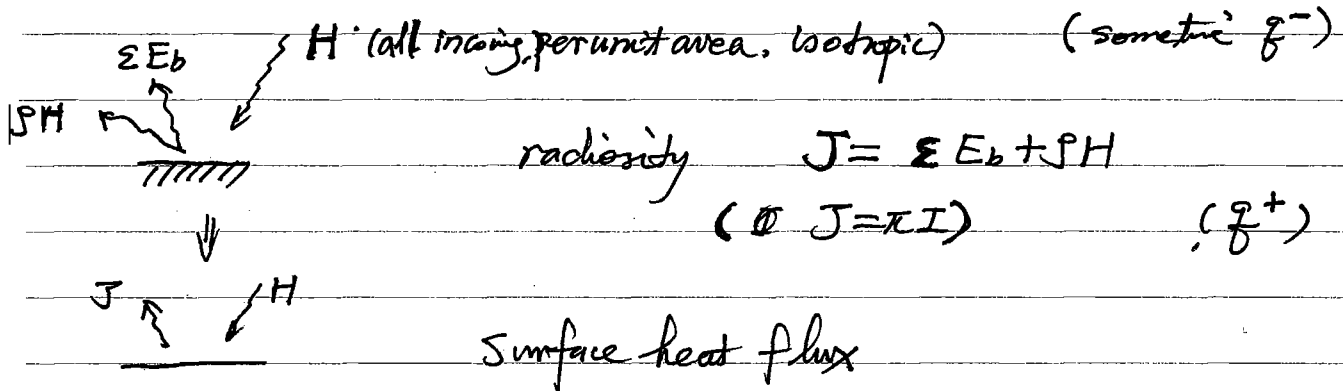
$$Q_{12} = Q_{1 \rightarrow 2} - Q_{2 \rightarrow 1}$$

$$= A_1 F_{12} \sigma T_1^4 - A_2 F_{21} \sigma T_2^4$$

$$= A_1 F_{12} \sigma (T_1^4 - T_2^4)$$

Sanity check  $T_1 = T_2 \quad Q_{12} = 0$

• Diffuse-gray surfaces: Radiosity & Irradiation



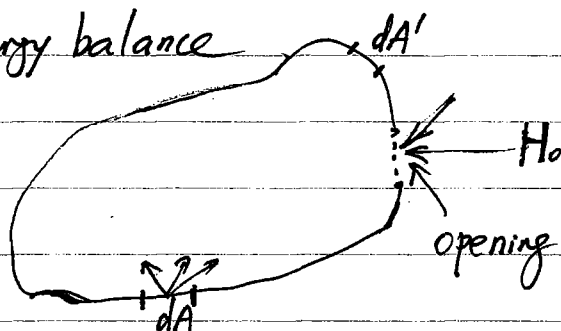
$$q = J - H = \epsilon E_b - \alpha H.$$

~~viewed from outside~~  $H = J q$   $\left\{ \begin{array}{l} \text{view from inside} \\ \text{viewed from outside} \end{array} \right.$

$$q = \epsilon E_b - \alpha J + \alpha q$$

$$q = \frac{\epsilon (E_b - J)}{1 - \alpha}$$

• Energy balance



← viewed from immediate inside

$$f(\vec{r}) = \epsilon(\vec{r}) E_b(\vec{r}) - \alpha(\vec{r}) H(\vec{r}) \quad \Rightarrow \quad J(\vec{r}) = \epsilon(\vec{r}) E_b(\vec{r}) + \rho(\vec{r}) H(\vec{r})$$

$$= \mathbf{J}(\vec{r}) - H(\vec{r})$$

↳ viewed from outside of ~~the~~ surface

$$f(\vec{r}) = \frac{\epsilon(\vec{r})}{1-\epsilon(\vec{r})} [E_b]$$

$$dA_{\vec{r}} H(\vec{r}) = \int J(\vec{r}') dF_{dA-dA'} \cdot dA' + H_0(\vec{r}) dA$$

$\downarrow$   
 $dF_{dA-dA'} \cdot dA$

↳ external arriving at  $dA$

⇒

$$H(\vec{r}) = \int_A J(\vec{r}') dF_{dA-dA'} + H_0(\vec{r})$$

$$f(\vec{r}) = \epsilon(\vec{r}) E_b(\vec{r}) - \alpha(\vec{r}) \left[ \int_A J(\vec{r}') dF_{dA-dA'} + H_0(\vec{r}) \right]$$

or

$$J(\vec{r}) = \epsilon(\vec{r}) E_b(\vec{r}) + \rho(\vec{r}) \left[ \int_A J(\vec{r}') dF_{dA-dA'} + H_0(\vec{r}) \right]$$

↑

↑ if  $T(\vec{r})$  is known.

Integral equation to solve for  $J(\vec{r})$

↳  $E_b(\vec{r})$  is unknown, but  $f(\vec{r})$  is known, substitute  $f = \frac{\epsilon(E_b - J)}{1-\epsilon}$

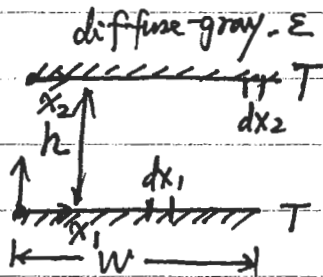
$$J(\vec{r}) = f(\vec{r}) + \int_A J(\vec{r}') dF_{dA-dA'} + H_0(\vec{r})$$

We can also eliminate  $J(\vec{r}) \Rightarrow$

$$\frac{f(\vec{r})}{\epsilon(\vec{r})} = \int_A \left( \frac{1}{\epsilon(\vec{r}')} - 1 \right) f(\vec{r}') dF_{dA-dA'} + H_0(\vec{r})$$

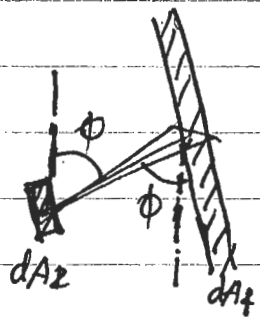
$$= E_b(\vec{r}) - \int_A E_b(\vec{r}') dF_{dA-dA'}$$

Example: local radiative flux  $q(x)$



$$J_2(x_2) = \epsilon \sigma T^4 + (1-\epsilon) \int_0^W J_1(x_1) dF_{d2-d1}$$

Assuming no radiation from ambient enters the gap (zero ambient).



$$\cos \phi = \frac{h}{\sqrt{h^2 + (x_2 - x_1)^2}}$$

$d\phi \rightarrow$  for  $dF_{d2-d1}$   $dx_1$  varying

$$dF_{d2-d1} = \frac{\cos \phi d\phi}{2}$$

~~$$d\phi = \frac{h}{2[h^2 + (x_2 - x_1)^2]^{3/2}} dx_1$$~~

$$d\phi = \frac{dx_1 \cos \phi}{\sqrt{h^2 + (x_2 - x_1)^2}}$$

$dx_1 \cos \phi$  projection of  $dx_1$  into  $\phi$  changing direction

$$J_2(x_2) = \epsilon \sigma T^4 + (1-\epsilon) \int_0^W \frac{h^2 J_1(x_1)}{2[h^2 + (x_2 - x_1)^2]^{3/2}} dx_1$$

Symmetry  $J_1(x_1) = J_2(x_2)$

$$J_2(x_2) = \epsilon \sigma T^4 + \frac{(1-\epsilon) h^2}{2} \int_0^W \frac{J_2(x_1)}{[h^2 + (x_2 - x_1)^2]^{3/2}} dx_1$$

$$\xi = \frac{x}{h}, \quad \bar{j} = \frac{J}{\sigma T^4}, \quad \bar{W} = w/h$$

$$\bar{j}(\xi) = \epsilon + \frac{1-\epsilon}{2} \int_0^{\bar{W}} \underbrace{\frac{1}{[1 + (\xi' - \xi)^2]^{3/2}}}_{\text{Kernel}} \bar{j}(\xi') d\xi'$$

Integral equation  ~~$\bar{j}(\xi)$~~

(1st kind  $\epsilon=0$ , no extn term)

Fredholm equation of the 2nd kind

If  $\epsilon_1 \neq \epsilon_2$  or  $T_1 \neq T_2$ ,  $\Rightarrow$  2 coupled integral equations.

Solution method for integral equations

- (1) Successive approximation
- (2) variational calculus
- (3) Kernel approximation
- (4) Numerical quadrature

Numerical quadrature

$$\int_a^b f(\xi, \xi') d\xi' \approx (b-a) \sum_{j=1}^N w_j f(\xi, \xi_j)$$

$\uparrow$   
 weight coefficient  
 (quadrature coefficient)

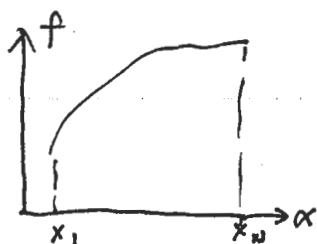
Apply to our equation

$$j(\xi) = \epsilon + \frac{1-\epsilon}{2} \sum_{j=1}^N w_j k(\xi, \xi_j) j(\xi_j)$$

Next set  $\xi = \xi_1, \xi_2, \dots, \xi_N$ .  $\Rightarrow$  a set of  $N$  equations to solve.  $\Rightarrow$  matrix inversion

Simplest.

Trapezoidal rule



$$\int_{x_1}^{x_N} f(x) dx = h \left[ \frac{1}{2} f_1 + f_2 + \dots + f_{N-1} + \frac{1}{2} f_N \right] + O(N^{-2})$$

Simpson rule  $- O(N^{-4})$

equally spaced points

Gaussian Quadrature: both weight  $W_j$  &  $x_j$   
good if  $f$  is well approximated by polynomial

$$\int_a^b W(x) f(x) dx \approx \sum_{j=1}^N W_j f(x_j)$$

↑  
known funct.

↑  
exact if  $f(x)$  is polynomial of order  $N$

Gauss-Legendre  $W=1, -1 \leq x \leq 1$

$$W_j = \frac{2}{(1-x_j^2) [P'_N(x_j)]^2}$$

$$(j+1) P_{j+1} = (2j+1)x P_j - j P_{j-1}$$

$P_j$  - Legendre polynomial

~~$P_j$~~

Kernel ill behaved:

at  $x=x', K \rightarrow \infty$

~~$$g(x) = f(x) + \int_a^b K(x, x') g(x') dx'$$~~

~~$$g(x) = f(x) + \int_a^b K(x, x') [g(x') - g(x)] dx'$$~~

$$+ g(x) \int_a^b K(x, x') dx'$$

↳ can be integrated analytically or numerically.

back to the problem

more heat loss at  
2 edges.

